

Radiative transfer : the diffusion approximation

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Radiation hydrodynamics

- ▶ **Supersonic, hypersonic flows**
 - Compression of matter
 - High temperature
 - Photon emission

- ▶ **Radiation : change the dynamics and morphology of flows**



Stellar jets ©NASA, ESA, & M. Livio



SNR ©Digitized sky Survey, ESA/ESO/NASA

HADES 2D

- ▶ Numerical code for radiation hydrodynamics models simulations
 - Finite volume approach
 - Parallelized code written in Fortran 90

- ▶ Computation of physical quantities as a function of time

Hydrodynamic quantities :

- Density : ρ
- Velocity : U
- Energy : E

Radiative quantities :

- Radiative energy : E_R
- Radiative flux F_R
- Radiative pressure : P_R

Radiation hydrodynamics equations

Euler equations with M1 multigroup model coupling :

$$\text{Coupling} \left\{ \begin{array}{l} \text{Hydro} \left\{ \begin{array}{l} \partial_t \rho + \nabla \cdot (\rho \mathbf{u}) = 0, \\ \partial_t (\rho \mathbf{u}) + \nabla \cdot (\rho (\mathbf{u} \otimes \mathbf{u}) + p \mathbb{I}) = \sum_{g=1}^{\mathcal{G}} \mathbf{S}_g, \\ \partial_t E + \nabla \cdot (\mathbf{u} (E + p)) = \sum_{g=1}^{\mathcal{G}} c S_g^0, \end{array} \right. \\ \\ \text{Rad} \left\{ \begin{array}{l} \partial_t E_{R_g} + \nabla \cdot \mathbf{F}_{R_g} = -c S_g^0, \\ \partial_t (c^{-2} \mathbf{F}_{R_g}) + \nabla \cdot \mathbf{P}_{R_g} = -\mathbf{S}_g, \end{array} \right. \quad g = 1, \dots, \mathcal{G}. \end{array} \right.$$

Turpault, PhD thesis 03 ; Lowrie et al., JQRST 01 ; Mihalas & Auer, JQRST 01

M1 multigroup model

Source terms (TLE) :

$$S^0 = \kappa_P (E_R - a_R T^4),$$

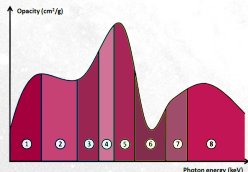
$$\mathbf{S} = \kappa_R \mathbf{F}_R / c,$$

with resp. κ_P and κ_R Planck and Rosseland mean opacities :

$$\kappa_P = \frac{\int_{\nu} \kappa(\nu) B(\nu, T) d\nu}{\int_{\nu} B(\nu, T) d\nu},$$

$$\kappa_R^{-1} = \frac{\int_{\nu} \chi^{-1}(\nu) \partial_T B(\nu, T) d\nu}{\int_{\nu} \partial_T B(\nu, T) d\nu}.$$

- Generally, **strong fluctuation** of opacities
- **Segmentation** of frequencies ν in \mathcal{G} groups
- Opacity calculations in HADES : use of tables of opacities which are results from atomic physics calculations. Interpolation on the grid (ρ, T) .



Diffusion approximation

λ : mean free path of a photon L : characteristic length of the phenomena

- ▶ **Optically very thick medium** : $\lambda \ll L$
 - Fluid opaque to photons, local radiative phenomena, diffusion approximation

Radiative quantities :

$$E_R = a_R T^4 \qquad F_R = -\frac{1}{3} \frac{c}{\kappa_R} \nabla E_R \qquad P_R = \frac{1}{3} E_R \mathbb{I}$$

Equation of state for perfect gas : $T = m(\gamma - 1)/k_B(E/\rho - \|\mathbf{u}\|^2/2)$

Equations in the diffusion approximation :

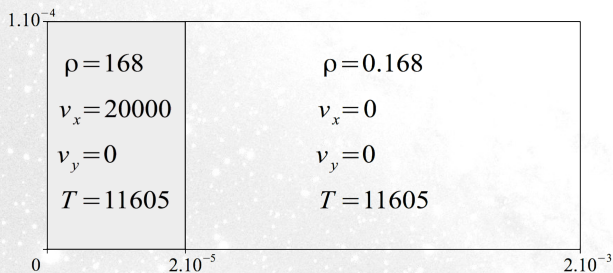
$$\partial_t \rho + \nabla \cdot (\rho \mathbf{u}) = 0,$$

$$\partial_t (\rho \mathbf{u}) + \nabla \cdot (\rho (\mathbf{u} \otimes \mathbf{u}) + p \mathbb{I}) = -\frac{1}{3} \nabla E_R,$$

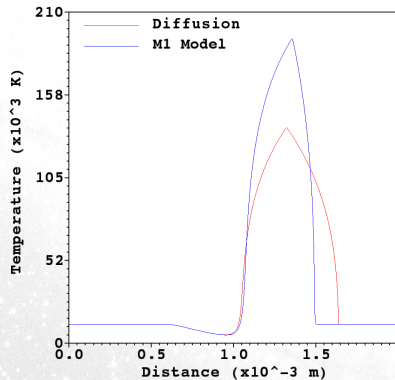
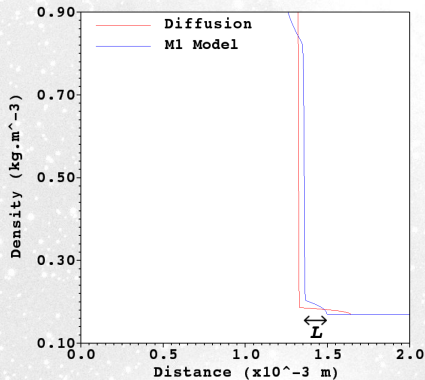
$$\partial_t E + \nabla \cdot (\mathbf{u} (E + p)) = -\partial_t E_R + \frac{1}{3} \frac{c}{\kappa_R} \Delta E_R - \frac{4}{3} \nabla \cdot (\mathbf{u} E_R).$$

Stationary radiative shock

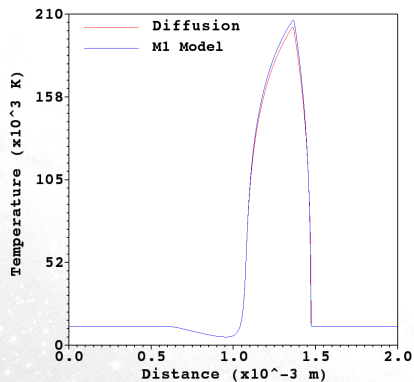
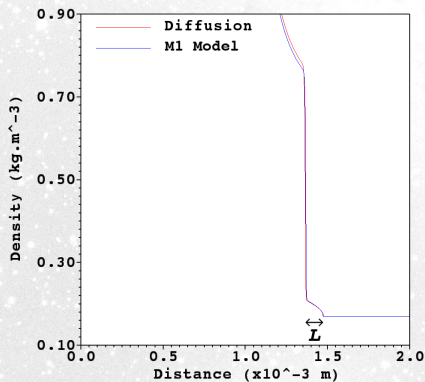
- Comparison between diffusion approximation and the M1 model according to the mean free path of photons :



- $M_x = 2000$, $M_y = 10$ and $t_f = 4.10^{-8}$ s
- Mach number $M \approx 8$

$\lambda = 10$ microns

$$\frac{\lambda}{L} \approx 0.07$$

$\lambda = 1$ micron

$$\frac{\lambda}{L} \approx 0.01$$

Computation time : M1 multigroup model : 12 h, diffusion approximation : 35 min

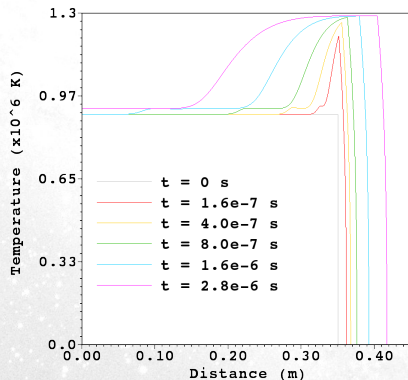
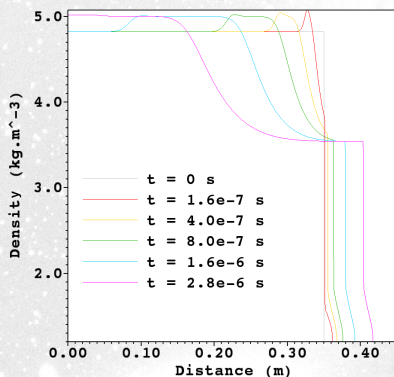
Stationary radiative shock in shock frame

- Instead λ constant, we chose $\sigma = 1/(\rho\lambda) = 4.10^3 \text{ m}^2.\text{kg}^{-1}$:

5.10^{-4}	$\rho_2 \approx 4.81$	$\rho_1 = 1.2$
	$v_{x2} \approx -50000$	$v_{x1} = -200000$
	$v_{y2} = 0$	$v_{y1} = 0$
	$T_2 \approx 900000$	$T_1 = 10$
-3.510^{-1}	0	1.10^{-1}

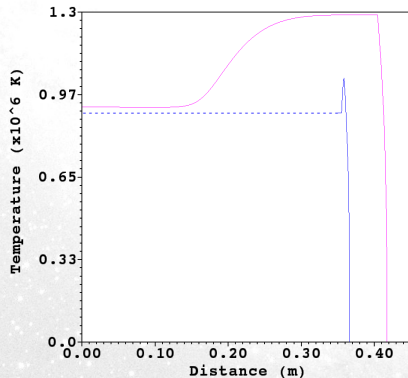
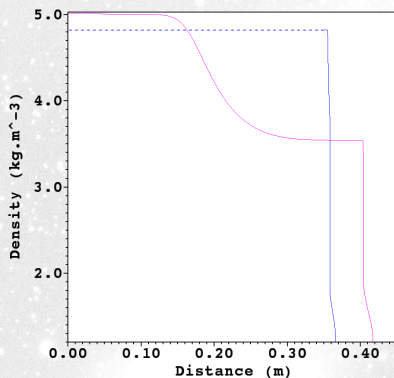
- $M_x = 4500$ and $M_y = 5$
- Mach number $M \approx 540$

Stationary radiative shock in shock frame



Stationary radiative shock in shock frame

- ▶ In the literature : computation are not lead in the post-choc medium (Dirichlet boundary condition)



- ▶ Peak for temperature

Conclusion

- Radiative transfer : expensive simulations
- Diffusion approximation : simpler description of transfer
- Lower computation times
- Useful to prepare simulations with the M1 multigroup model

Thank you for your attention !