## "Relativity & Compact Objects" team: 2014-2015 report

### Jérôme Novak, on behalf of the ROC team

Laboratoire Univers et Théories (LUTH) CNRS / Observatoire de Paris / Université Paris-Diderot

Journée du LUTH 2016, January, 20<sup>th</sup> 2016

## WHAT IS ROC?





Originally founded by Silvano Bonazzola & Jean-Alain Marck, with interests in:

- relativistic astrophysics,
- gravitational waves,
- numerical relativity,
- it has grown and evolved with time.

 $\Rightarrow$  8 permanent researchers, 1 with fixed-term contract & 4 PhD students.

Newer research directions:

- solutions in classical field theory,
- properties of nuclear matter,
- alternative theories of gravity,

• . . .

## IN THE LAST COUPLE OF YEARS...



- Microphysics for core-collapse supernovae
- Pulsars and neutron stars
- Hyperons in compact stars
- Gravitational waves
- Compact object binaries
- Rotating boson stars
- Scalar breathers
- SageManifolds *et al.*

## CORE-COLLAPSE SUPERNOVAE S. Bonazzola, I. Cordero-Carrión, J. Novak,

M. Oertel & collaborators

Study of local properties of matter for gravitational collapse of massive stars ( $\gtrsim 10 M_{\odot}$ ):

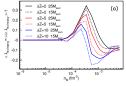
 $p^+ + e^- \to n + \nu_e$ 

• development of neutrino transport using spectral methods: GR effects taken into account ⇒first proof of principle, able to run on single-CPU in 6D (low resolution)



Peres et al. (2014)

 study of the influence of the presence of more neutrons in nuclei onto the electron capture rates in core-collapse. Modification of nuclear structure and up to 30% of EC rate ⇒quite important impact on dynamics



Raduta et al. (2015)

## HYPERONS IN NUCLEAR MATTER

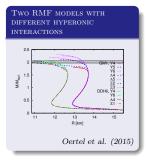
M. Oertel & collaborators

Nuclear physics models prediction:

- hyperons appear at  $n_B \sim 2 3n_0$ ,
- maximum neutron star masses of  $\sim 1.4 M_{\odot}$

 $\Rightarrow$ need short-range repulsion to stiffen the equation of state:

- With quark matter appearing early (very early!) enough,
- Modify the interaction
  - In microscopic models (BHF) this seems to be a problem
  - In phenomenological models not difficult Here: different RMF models for NS with hyperons
    - Maximum masses above  $2M_{\odot}$
    - Large range of radii for intermediate mass stars
    - Considerable hyperon fraction in NS

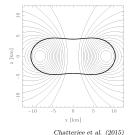


## MAGNETIC FIELD IN NEUTRON STARS & AROUND PULSARS S. Bonazzola, D. Chatterjee, E. Gourgoulhon, J. Novak, M. Oertel & collaborators

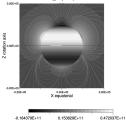
Magnetic field

New numerical models with magnetic field in neutron stars:

- Mixed poloidal-toroidal magnetic field configurations: formalism & first numerical application in Uryū *et al.* (2014).
- Inclusion of magnetization and magnetic field-dependent equation of state in global models by Chatterjee *et al.* (2015).







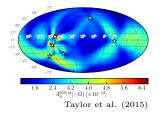
Generalised multipole solution for a rotating star with a magnetic field (Bonazzola *et al.* 2015):

- multipole expansion of electric and magnetic fields around a rotating star (without plasma)
- Explicit form of the solution ⇒current closure in pulsar magnetospheres ?

Bonazzola et al. (2015)

## PULSAR TIMING ARRAY G. THEUREAU & COLLABORATORS

Limits on supermassive black hole binary gravitational wave background from European Pulsar Timing Array



- millisecond pulsars used as probes (clocks) for the detection of nanohertz gravitational waves
- mostly looking at supermassive black hole binaries, which may be few and give anisotropic background

 $\Rightarrow$  study with more general techniques to take into account possible anisotropy.

 $\Rightarrow$ strain amplitude for  $\ell > 0$  (spherical harmonics) lower than 40% of the monopole value.

Combination with galaxy catalogues?

## Compact object binaries

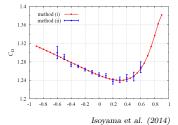
### A. Le Tiec & collaborators

Computations of gravitational waveforms need high-accuracy description of the binary orbits.

- Comparison between post-Newtonian and self-force approximations for binary black hole problem.
- Further link perturbative approach / numerical results: unexpectedly good agreement for equal-mass binaries.
- "First law of mechanics" for black hole binaries: fundamental relations between physical quantities allowing for better analytic model (EOB).

### EXAMPLE:

Computation of gravitational self-force for a particle around a Kerr black hole  $\Rightarrow$  shift in the frequency of the innermost stable circular orbit.



## ROTATING BOSON STARS

É. Gourgoulhon, Ph. Grandclément, C. Somé & collaborators

Self-gravitating complex scalar field in general relativity:

- Stationarity and axisymmetry of the spacetime.
- Ansatz on the scalar field:  $\Phi = \phi \left( r, \theta \right) \exp \left[ i \left( \omega t - k \varphi \right) \right].$
- Choice of scalar field potential  $V\left(|\Phi|^2\right)$ : free field, additional repulsive term or solitonic.

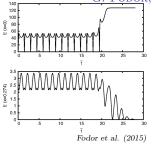
 $\Rightarrow$ alternative model to black hole for the central region of our Galaxy (tests with GRAVITY);  $\Rightarrow$ computation of accretion torus images, to see possible observable differences with black holes



Vincent et al. (2016)

## Scalar breathers

G. Fodor, Ph. Grandclément & collaborators



Scalar field soliton-like configuration in anti-de Sitter (AdS) spacetime or self-gravitating asymptotically AdS in D-dimensions.

 $\Rightarrow$ spatially localized, spherically symmetric & time periodic solutions

### "SELF-GRAVITATING SCALAR BREATHERS"

- KADATH library: spectral methods both in space and time (Fourier transform), to build initial models
- discrete family of solutions, labelled by their frequency
- time evolution to look for stability ⇒unstable branch beyond critical central density

# SAGEMANIFOLDS

É. GOURGOULHON & COLLABORATORS

- Package used with the mathematics software SageMath
- Implements differential geometry and tensor calculus on real differential manifolds of arbitrary dimension.

| $ \begin{array}{l} \cdot \cdot$   | $\frac{a^{5}q^{2}+6a^{3}n}{\nabla^{6}\mathbf{x}^{p}_{k}\mathbf{R}^{5}\mathbf{q}^{2}+\mathbf{\nabla}^{6}}$ |
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- Deals with tensor fields and not tensor components: various charts can be used, with corresponding representations of the tensor
- Written in Python under GNU Public License

 $\Rightarrow$  currently being integrated into SageMath and tensor calculus made parallel (M. Mancini)

 $\Rightarrow$ new version 0.9, see http://sagemanifolds.obspm.fr

## OPEN NUMERICAL TOOLS

### EVERYBOBY?

Longstanding tradition of opening our numerical projects to the whole community (and beyond):

- LORENE: solving Einstein equation, spectral methods & spherical coordinates; open source and freely downloadable since 2001.
- Kadath: inspired from LORENE, more flexible in geometries and physical problems; able to run on 100's of CPUs.
- Gyoto: orbits and ray-tracing computations in analytic or numerical GR spacetimes (F.H. Vincent, T. Paumard, LESIA).

### CoCoNuT

General relativistic code for core-collapse supernova simulations ⇒parallel version (thanks to that of LORENE) running up to 512 CPUs (F. Roy, P. Cerdá-Durán & J. Novak).

### Compose

Online database with equations of state tables for astrophysics, nuclear physics, ... ⇒operational and publicly accessible ⇒European effort within (New)CompStar frameworks (J.-Y. Giot, T. Klähn, M. Oertel & S. Typel)

All these servers are maintained by LUTH ... Strong support from the LUTH computer team

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## WHAT'S NEXT?

### Young researchers

- new member: Jean-Philippe Bruneton ⇒alternative theories of gravity and alternative models to black holes.
- Miguel Marques  $\Rightarrow$ rotating, hot neutron star models with realistic equation of state
- Grégoire Martinon ⇒numerical models of self-confined gravitational waves (geons)... Models for dark matter?
- Aurélien Sourie ⇒glitches from superfluid models & gravitational waves from oscillating neutron stars.
- Daniela Pérez ⇒black hole mimickers, avoiding the formation of a singularity.

### COLLABORATIONS WITH OTHER TEAMS

• Zakaria Meliani ⇒accretion disks and tori around boson stars.

• . . .



#### GIGONDAS Appellation Gigondas Contrôlée

14% VOL.

750 ML

Mis en bouteilles au Domaine par Earl Gras Edmond et Fils Yves Gras et Rémy Pédréno, 84190 Gigondas, Vaucluse.

Thank you!